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## Classe Scienze Fisiche Matematiche Naturali RENDICONTI

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## Dual cross-sectional measures

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# RENDICONTI 

DELLE SEDUTE

# DELLA ACCADEMIA NAZIONALE DEI LINCEI Classe di Scienze fisiche, matematiche e naturali 

Seduta dell'II gennaio 1975<br>Presiede il Presidente della Classe Beniamino Segre

## SEZIONE I

(Matematica, meccanica, astronomia, geodesia e geofisica)

Geometria. - Dual cross-sectional measures. Nota di Erwin Lutwak, presentata ${ }^{\left({ }^{( }\right)}$dal Socio B. Segre.


#### Abstract

RiASSUnto. - Si definiscono «quermass integrali» duali a mezzo di formole di Kubota duali. Si esaminano le relazioni fra questi integrali ed i funzionali di Minkowski; in particolare, si trovano le disuguaglianze che corrispondono per dualità a quelle classiche.


The integral formulas of Kubota [3] permit a simple recursive definition of the $n+\mathrm{I}$ cross-sectional measures $\mathrm{W}_{0}, \mathrm{~W}_{1}, \cdots, \mathrm{~W}_{n}$ in Euclidean $n$-space, $\mathrm{R}^{n}$. In a one dimensional space, for convex compact $A, W_{0}(A)$ is defined to be the length of $A$ while $W_{1}(A)$ is defined to be 2. After defining the cross-sectional measures in ( $n$ - I)-space they are defined in Euclidean $n$-space by letting $\mathrm{W}_{0}(\mathrm{~A})$ equal $\mathrm{V}(\mathrm{A})$, the $n$-dimensional volume of A , and letting

$$
\begin{equation*}
\mathrm{W}_{i}(\mathrm{~A})=\frac{\mathrm{I}}{n \omega_{n-1}} \int_{\Omega} \mathrm{W}_{i-1}^{\prime}\left(\mathrm{A} \mid \mathrm{P}_{u}\right) \mathrm{dS}(u) \quad[i>0] \tag{I}
\end{equation*}
$$

In this integral $\Omega$ is the surface of the unit ball $U$ in $R^{n}$ while dS denotes the area element on $\Omega$. A $\mid \mathrm{P}_{u}$ is the projection of A onto the hyperplane $\mathrm{P}_{u}$ which is perpendicular to $u \in \Omega$ and passes through the origin. $\mathrm{W}_{i-1}^{\prime}$ denotes the ( $i$ I )-th cross-sectional measure in ( $n-1$ )-space while $\omega_{n-1}$ denotes the volume of the unit ball in $\mathrm{R}^{n-1}$.

The dual cross-sectional measures of the title refer to the measures obtained when the definition of the cross-sectional measures is altered by replacing $A \mid P_{u}$ in (I) by its dual $A \cap P_{u}$. These measures arise naturally in the examination of a radial addition that is analogous to Minkowski addi-
(*) Nella seduta dell'ı gennaio 1975 .

1.     - RENDICONTI 1975, Vol. LVIII, fasc. 1.
tion. They appear as coefficients in a polynomial expression analogous to the Steiner polynomial expression for the volume of a parallel body [3]. In addition, the natural extension of these measures can be used to examine a harmonic addition previously considered by Firey [I, 2], Steinhardt [7, p. 15] and Rockafellar [6, p. 21].

The setting for this paper will be $\mathrm{R}^{n}$. Compact convex sets with nonempty interiors will be called convex bodies. The space of all convex bodies which contain the origin in their interior, endowed with the Hausdorff topology, will be denoted by $K_{n}$. For a convex body A, we use $\mathrm{H}_{\mathrm{A}}$ and $\mathrm{F}_{\mathrm{A}}$ to denote its support and distance function, respectively. A compact set A is called a star body with respect to a point $a$ if for every $c$ in A the line $\{t a+(\mathrm{I}-t) c \mid 0<t \leq \mathrm{I}\}$ lies in the interior of A . The set of all star bodies with respect to the origin will be denoted by $S_{n}$. Associated with a star body $\mathrm{A} \in S_{n}$ is a radial function $\rho_{\mathrm{A}}$ defined on $\Omega$ by:

$$
\rho_{\mathrm{A}}(u)=\operatorname{Sup}\{\lambda>0 \mid \lambda u \in \mathrm{~A}\} \quad[u \in \Omega]
$$

The radial function of a star body in $S_{n}$ is a positive, continuous real-valued function on $\Omega$. Conversely, a positive, continuous real-valued function on $\Omega$ is the radial function of a unique star body in $S_{n}$. A metric d can be defined in $S_{n}$ by letting

$$
\mathrm{d}(\mathrm{~A}, \mathrm{~B})=\operatorname{Sup}_{u \in \Omega}\left|\rho_{\mathrm{A}}(u)-\rho_{\mathrm{B}}(u)\right| \quad\left[\mathrm{A}, \mathrm{~B} \in S_{n}\right]
$$

It is easy to verify that $S_{n}$ endowed with the topology induced by this metric has $K_{n}$ as a subspace.

While the cross-sectional measures are defined for compact convex sets, the dual cross-sectional measures $\stackrel{W}{W}_{0}, \stackrel{W}{W}_{1}, \cdots, \widetilde{W}_{n}$ will be defined for star bodies in $S_{n}$.

Definition i. In $\mathrm{R}^{1}$ the dual cross-sectional measures are defined by

$$
\stackrel{\mathrm{W}}{0}^{(\mathrm{A})=\mathrm{V}(\mathrm{~A}) \quad \stackrel{\mathrm{W}}{1}(\mathrm{~A})=\omega_{1} \quad\left[\mathrm{~A} \in S_{1}\right] . . . ~}
$$

After defining the dual cross-sectional measures in Euclidean ( $n-1$ )-space they are defined in $\mathrm{R}^{n}$ by letting $\stackrel{W}{\mathrm{~W}}_{0}(\mathrm{~A})=\mathrm{V}(\mathrm{A})$ and

$$
\stackrel{\mathrm{W}}{i}(\mathrm{~A})=\frac{\mathrm{I}}{n \omega_{n-1}} \int_{\Omega} \check{\mathrm{W}}_{i-1}^{\prime}\left(\mathrm{A} \cap \mathrm{P}_{u}\right) \mathrm{dS}(u) \quad\left[\begin{array}{ll}
i>0 & \mathrm{~A} \in S_{n} \tag{2}
\end{array}\right]
$$

where $\tilde{\mathrm{W}}_{i-1}^{\prime}$ denotes the ( $i$ - )-th dual cross-sectional measure in Euclidean ( $n$ - I)-space.

Comparing (2) with (I), we obtain:
Theorem i.

$$
\stackrel{\rightharpoonup}{\mathrm{W}}_{i}(\mathrm{~A}) \leq \mathrm{W}_{i}(\mathrm{~A}) \quad\left[0<i<n \quad \mathrm{~A} \in K_{n}\right]
$$

with equality if and only if A is an $n$-ball (centered at the origin).

The $i$-th dual cross-sectional measure $\stackrel{\rightharpoonup}{W}_{i}$ is a map

$$
\stackrel{\rightharpoonup}{\mathrm{W}}_{i}: \quad S_{n} \rightarrow \mathrm{R} .
$$

It is continuous, bounded, additive, positive, rotation invariant, homogeneous of degree $n-i$ and monotone under set inclusion. All of these properties are simple consequences of our next theorem which describes the dual crosssectional measures of a star body as means of powers of its radial function.

Theorem 2.

$$
\tilde{\mathrm{W}}_{i}(\mathrm{~A})=\frac{\mathrm{I}}{n} \int_{\Omega} \rho_{\mathrm{A}}^{n-i}(u) \mathrm{dS}(u) \quad\left[\mathrm{A} \in S_{n}\right]
$$

The proof follows by induction on the dimension of the space using standard techniques (see Hadwiger [3, p. 212]).

The cross-sectional measures satisfy the cyclic inequality [3, p. 282]:

$$
\mathrm{W}_{j}^{k-i}(\mathrm{~A}) \geq \mathrm{W}_{i}^{k-j}(\mathrm{~A}) \mathrm{W}_{k}^{j-i}(\mathrm{~A}) \quad\left[i<j<k \quad \mathrm{~A} \in K_{n}\right] .
$$

As a simple consequence of Hölder's Inequality [4, p. I40] we have:
Theorem 3.

$$
\left.\stackrel{\stackrel{W}{\mathrm{~W}}}{j}_{k-i}^{(\mathrm{A}) \leq \stackrel{\rightharpoonup}{\mathrm{W}}_{i}^{k-j}(\mathrm{~A}) \stackrel{\mathrm{W}}{k}_{j-i}(\mathrm{~A}) \quad[i<j<k} \quad \mathrm{A} \in S_{n}\right]
$$

with equality if and only if A is an $n$-ball (centered at the origin).
We note, that for convex bodies in $K_{n}$, Theorem 3 is a consequence of a general inequality between dual mixed volumes that was obtained by us in [5].

The Minkowski sum A $+B$ of two convex bodies A and B can be defined by the equation

$$
\mathrm{H}_{\mathrm{A}+\mathrm{B}}=\mathrm{H}_{\mathrm{A}}+\mathrm{H}_{\mathrm{B}} .
$$

Given two star bodies $\mathrm{A}, \mathrm{B} \in S_{n}$ we define the radial sum $\mathrm{A} \otimes \mathrm{B}$ by:
Definition 2.

$$
\rho_{\mathrm{A} \otimes \mathrm{~B}}=\rho_{\mathrm{A}}+\rho_{\mathrm{B}} \quad\left[\mathrm{~A}, \mathrm{~B} \in S_{n}\right] .
$$

The Brunn-Minkowski Theorem [3, p. 187] states that:

$$
\mathrm{V}^{1 / n}(\mathrm{~A}+\mathrm{B}) \geq \mathrm{V}^{1 / n}(\mathrm{~A})+\mathrm{V}^{1 / n}(\mathrm{~B}) \quad\left[\mathrm{A}, \mathrm{~B} \in K_{n}\right]
$$

with equality if and only if A and B are homothetic. A simple application of the Minkowski Inequality [4, p. 146] yields:

Theorem 4.

$$
\mathrm{V}^{1 / n}(\mathrm{~A} \otimes \mathrm{~B}) \leq \mathrm{V}^{1 / n}(\mathrm{~A})+\mathrm{V}^{1 / n}(\mathrm{~B}) \quad\left[\mathrm{A}, \mathrm{~B} \in S_{n}\right]
$$

with equality if and only if A is a dilation of B (with the origin as the center of dilation).

For a convex body A and a scalar $\mu>0$ the parallel body $\mathrm{A}_{\mu}$ is defined to be $\mathrm{A}+\mu \mathrm{U}$. For a star body $\mathrm{A} \in S_{n}$ and $\mu>0$ we define the radial body $A$ by:

Definition 3.

$$
{ }_{\mu} \mathrm{A}=\mathrm{A} \otimes \mu \mathrm{U} \quad\left[\mu>\mathrm{O} \quad \mathrm{~A} \in S_{n}\right] .
$$

For the volume of the parallei body $\mathrm{A}_{\mu}$ we have the Steiner polynomial expression [3, p. 214]:

$$
\mathrm{V}\left(\mathrm{~A}_{\mu}\right)=\sum_{i=0}^{n}\binom{n}{i} \mathrm{~W}_{i}(\mathrm{~A}) \mu^{i} \quad\left[\mu>0 \quad \mathrm{~A} \in K_{n}\right]
$$

Just as the cross-sectional measures appear as coefficients in the polynomial expression of $\mathrm{V}\left(\mathrm{A}_{\mu}\right)$, the dual cross-sectionai measures appear as coefficients in a polynomial expression of $V(\mu \mathrm{~A})$.

Theorem 5.

$$
\mathrm{V}(\mu \mathrm{~A})=\sum_{i=0}^{n}\binom{n}{i} \stackrel{\mathrm{~W}}{i}(\mathrm{~A}) \mu_{0}^{i} \quad\left[\mu>0 \quad \mathrm{~A} \in S_{n}\right] .
$$

To prove this we merely note that $\rho_{\mu \mathrm{A}}=\rho_{\mathrm{A}}+\mu$.
Combining the definition of the surface area of a convex body [3, p. 184] with the Cauchy area formula [3, p. 208] we obtain:

$$
\operatorname{Lim}_{\mu \rightarrow 0}\left[\mathrm{~V}\left(\mathrm{~A}_{\mu}\right)-\mathrm{V}(\mathrm{~A})\right] / \mu=\frac{\mathrm{I}}{\omega_{n-1}} \int_{\Omega} \mathrm{V}^{\prime}\left(\mathrm{A} \mid \mathrm{P}_{u}\right) \mathrm{dS}(u) \quad\left[\mathrm{A} \in K_{n}\right]
$$

where $\mathrm{V}^{\prime}$ denotes the volume in Euclidean ( $n-\mathrm{I}$ )-space. As a direct consequence of Theorem 5 we have:

Corollary.

$$
\operatorname{Lim}_{\mu \rightarrow 0}[\mathrm{~V}(\mu \mathrm{~A})-\mathrm{V}(\mathrm{~A})] / \mu=\frac{\mathrm{I}}{\omega_{n-1}} \int_{\Omega} \mathrm{V}^{\prime}\left(\mathrm{A} \cap \mathrm{P}_{n}\right) \mathrm{dS}(u) \quad\left[\mathrm{A} \in S_{n}\right]
$$

As presented in Definition I the dual cross-sectional measures $\stackrel{W}{W}_{i}$ have indices $i$ restricted to integer values between zero and $n$. However, Theorem I points to a natural extension of the definition so that the $\stackrel{W}{W}_{i}$ are defined for all real indices.

Definition $\mathrm{I}^{*}$.

$$
\stackrel{\mathrm{W}}{i}^{(\mathrm{A})=} \int_{\Omega} \rho_{\mathrm{A}}^{n-i}(u) \mathrm{dS}(u) \quad\left[i \in \mathrm{R} \quad \mathrm{~A} \in S_{n}\right]
$$

The new $\stackrel{\rightharpoonup}{\mathrm{W}}_{i}$ are also positive, continuous, additive, rotation invariant and homogeneous of degree $n-i$. However, they are bounded and monotone only for $i \leq n$. Theorem 3 remains unaltered if we allow the indices of the dual cross-sectional measures to range over all real numbers.

With extended indices the dual cross-sectional measures can be used to examine a harmonic addition considered by Firey [I, 2], Steinhardt [7, p. I5] and Rockafellar [6, p. 2I].

The harmonic sum $\mathrm{A} \times \mathrm{B}$ of two convex bodies $\mathrm{A}, \mathrm{B} \in K_{n}$ is defined by:
Definition 4.

$$
\mathrm{F}_{\mathrm{A} \times \mathrm{B}}=\mathrm{F}_{\mathrm{A}}+\mathrm{F}_{\mathrm{B}} \quad\left[\mathrm{~A}, \mathrm{~B} \in K_{n}\right] .
$$

We note that, while we chose not to do so, harmonic addition could have been defined in $S_{n}$ by letting $f_{A \times B}=\left(\rho_{A}^{-1}+\rho_{B}^{-1}\right)^{-1}$. Both definitions coincide in $K_{n}$.

The following dual to the Brunn-Minkowski theorem is due to Firey [r] and Steinhardt [7]:

Theorem 6.

$$
\mathrm{V}^{-1 / n}(\mathrm{~A} \times \mathrm{B}) \geq \mathrm{V}^{-1 / n}(\mathrm{~A})+\mathrm{V}^{-1 / n}(\mathrm{~B}) \quad\left[\mathrm{A}, \mathrm{~B} \in K_{n}\right]
$$

with equality if and only if A is a dilation of B (with the origin as the center of dilation).

The scalar product $\mu \mathrm{A}$ of a convex body A and a scalar $\mu>0$ can be defined by the equation $\mathrm{H}_{\mu \mathrm{A}}=\mu \mathrm{H}_{\mathrm{A}}$. Analagously, we define a harmonic scalar product $\mu \circ \mathrm{A}$ by:

Definition 5.

$$
\mathrm{F}_{\mu_{\circ} \mathrm{A}}=\mu \mathrm{F}_{\mathrm{A}} \quad\left[\mu>0 \quad \mathrm{~A} \in K_{n}\right] .
$$

We note that harmonic scalar products could have been defined for star bodies A $\in S_{n}$ by letting $\rho_{\mu_{0} A}=\mu^{-1} \rho_{\mathrm{A}}$. Both definitions coincide in $K_{n}$.

Analogous to the definition of the parallel body we define the harmonic body $\mathrm{A}^{\mu}$ by:

Definition 6.

$$
\mathrm{A}^{\mu}=\mathrm{A} \times \mu \circ \mathrm{U} \quad\left[\mu>0 \quad \mathrm{~A} \in K_{n}\right] .
$$

Our last theorem shows that, for small $\mu$, the extended dual cross-sectional measures appear as coefficients in an expression for $V\left(A^{\mu}\right)$.

Theorem 7.

$$
\mathrm{V}\left(\mathrm{~A}^{\mu}\right)=\sum_{i=0}^{\infty}\binom{-n}{i} \stackrel{\mathrm{~W}}{-i}(\mathrm{~A}) \mu^{i} \quad\left[\mu<\operatorname{Inf} \rho_{\mathrm{A}}\right]
$$

The proof is a simple exercise involving the use of the binomial theorem.

## References

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