## ATTI ACCADEMIA NAZIONALE DEI LINCEI

# CLASSE SCIENZE FISICHE MATEMATICHE NATURALI

# RENDICONTI

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# Characterizations of Einstein Kaehler manifolds and applications

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Geometria differenziale. — Characterizations of Einstein Kaehler manifolds and applications. Nota di Bang-yen Chen (\*), presentata (\*\*) dal Socio B. Segre.

RIASSUNTO. — Vengono date condizioni sufficienti affinché una varietà compatta di Kaehler o coomologicamente di Einstein-Kaehler sia einsteiniana (Teorema 1, 2); se ne deducono condizioni assicuranti che un'intersezione completa in uno spazio proiettivo complesso risulti uno spazio lineare od un'iperquadrica (Teorema 3).

### I. STATEMENT OF RESULTS

Let M be an *n*-dimensional compact Kaehler manifold. Let  $\theta^1, \dots, \theta^n$  be a local field of unitary coframes with the Kaehler metric g and the Ricci tensor S given by

$$\begin{split} g &= \tfrac{1}{2} \, \Sigma \, (\theta^i {\otimes} \bar{\theta}^i + \bar{\theta}^i {\otimes} \theta^i) \,, \\ \mathrm{S} &= \tfrac{1}{2} \, \Sigma \, (\mathrm{R}_{i \overline{\jmath}} \, \theta^i {\otimes} \bar{\theta}^j + \bar{\mathrm{R}}_{i \overline{\jmath}} \, \bar{\theta}^i {\otimes} \theta^j) \,, \end{split}$$

respectively. The fundamental 2-form  $\Phi$  and the Ricci form  $\gamma$  are then given respectively by

$$\Phi = \frac{\sqrt{-1}}{2} \, \Sigma \theta^i \wedge \bar{\theta}^i$$

(2) 
$$\gamma = \frac{\sqrt{-1}}{4\pi} \sum_{i,j} \theta^{i} \wedge \bar{\theta}^{j}.$$

Let  $[\sigma]$  denote the cohomology class represented by  $\sigma$ . It is well known that the first Chern class  $c_1$  of M is represented by  $\gamma$  and the last de Rham cohomology group  $H^{2n}(M; R)$  is generated by  $[\Phi^n]$ .

We put

$$\omega = [\Phi]$$

and

$$\omega^{n-k} c_1^k = a_k \omega^n, \qquad k = 0, 1, \dots, n,$$

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where  $\omega^{n-k} c_1^k$  denotes the cup product of  $\omega^{n-k}$  and  $c_1^k$ . We define the *k*-th scalar curvature  $\rho_k$  by

$$\det \left( \delta_{ij} + t \mathbf{R}_{i\bar{j}} \right) = \sum_{k=0}^{n} \binom{n}{k} \, \rho_k \, t^k,$$

where  $\binom{n}{k}$  is the binomial coefficient. It is clear that  $\rho_0 = 1$ ,  $\rho_1$  is the (normalized) scalar curvature and  $\rho_n = \det(\mathbf{R}_{ij})$ .

We shall prove the following.

THEOREM 1. Let M be an n-dimensional compact Kaehler manifold. If

- (i)  $a_k^2 \le a_{k-1} a_{k+1}$ ;
- (ii)  $\rho_k$ ,  $\rho_{k+1}$  (or  $\rho_{k-1}$ ,  $\rho_k$ ) are positive constants,

for some k;  $1 \le k \le n-1$ , then M is Einsteinian, where  $a_i$ ; 1 = k-1, k, k+1, are given by  $\omega^{n-i} c_1^i = a_i \omega^n$ .

We say that M is cohomologically Einsteinian if  $c_1 = b\omega$  for some constant b. As applications of Theorem 1 we shall prove the following [2].

THEOREM 2. Let M be an n-dimensional compact cohomological Einstein Kaehler manifold. If there exists k,  $1 \le k \le n$ , such that  $\rho_{k-1}$  and  $\rho_k$  are positive constants, then M is Einsteinian.

Let  $P_{n+p}(C)$  be an (n+p)-dimensional complex projective space with the Fubini-Study metric of constant holomorphic sectional curvature 1. An n-dimensional algebraic submanifold in  $P_{n+p}(C)$  is called a *complete intersection* if M is given as an intersection of p nonsingular hypersurfaces of  $P_{n+p}(C)$  in general position. A complete intersection is always a Kaehler manifold by considering the induced Kaehler metric from  $P_{n+p}(C)$ .

Theorem 3. Let M be an n-dimensional complete intersection in  $P_{n+p}(C)$ . If there exists k,  $1 \le k \le n$ , such that  $\rho_{k-1}$  and  $\rho_k$  are positive constants, then M is either a linear subspace or a hyperquadric in some (n+1)-dimensional linear subspace.

Remark I. Assumptions (i) and (ii) are essential. For examples: (a) Let  $M = P_k(C) \times T^{n-k}$ , where  $T^{n-k}$  denotes an (n-k)-dimensional complex torus with the flat metric. Then  $\rho_{k+1} = 0$ ,  $\rho_k$  is constant and  $a_k^2 > a_{k-1} a_{k+1} = 0$ . (b) Let M be a algebraic hypersurface of  $P_{n+1}(C)$  with degree  $\neq I$ , 2. Then M is cohomologically Einsteinian (see the proof of Theorem 3), in particular, we have  $a_k^2 = a_{k-1} a_{k+1}$ , but  $\rho_k$ ,  $\rho_{k+1}$  are not constant, simultaneously.

Remark 2. If k = 1, the assumption of the constancy of  $\rho_{k-1}$  is automatically satisfied. In this case, Theorem 2 and Theorem 3 reduce to results of Kobayashi [5] and Hano [3], respectively.

Remark 3. For hypersurfaces see [1].

### 2. PROOF OF THEOREMS I AND 2

First we prove the following general Lemma.

LEMMA 1. Let M be an n-dimensional compact Kaehler manifold. Then

$$\int\limits_{\mathbf{M}} \rho_k * \mathbf{I} = (2 \pi)^k a_k \int\limits_{\mathbf{M}} * \mathbf{I},$$

where  $\omega^{n-k} c_1^k = a_k \omega^n$  and \* denotes the Hodge star operator.

*Proof.* Since  $\omega^{n-k} c_1^k = a_k \omega^n$ , there exists a (2 n - 1)-form  $\eta$  such that  $\Phi^{n-k} \wedge \gamma^k = a_k \Phi^n + d\eta.$ 

From the following identities:

$$*(\Phi^{n-k}\wedge\gamma^k)=\frac{n!\;\rho_k}{(2\;\pi)^k},\qquad k=0\;,\;1\;,\cdots,n\;,$$

we find

$$\rho_k = (2 \pi)^k a_k + \frac{(2 \pi)^k}{n!} * d\eta.$$

Thus by taking integration of both sides of this equation over M and by using the identity  $(* d\eta) * I = d\eta$ , we get the lemma.

Now we return to the proof of Theorem 1.

From assumption (ii)  $\rho_k$  and  $\rho_{k+1}$  (or  $\rho_{k-1}$  and  $\rho_k$ ) are constant, then from Lemma 1 we find

$$\int\limits_{M} \rho_{k}^{2} * I = (2 \pi)^{k} a_{k} \int\limits_{M} \rho_{k} * I = (2 \pi)^{2k} a_{k}^{2} \int\limits_{M} * I,$$

$$\int\limits_{M} \rho_{k-1} \rho_{k+1} * I = (2 \pi)^{2k} a_{k-1} a_{k+1} \int\limits_{M} * I.$$

Then by assumption (i) we find

(5) 
$$\int_{\mathbb{N}} (\rho_k^2 - \rho_{k-1} \rho_{k+1}) * I \leq 0.$$

On the other hand, from the definition of  $\rho_k$  and a well-known inequality on elementary symmetric functions we have

where the equality holds if and only if  $(R_{ij})$  is proportional to the identity matrix, i.e., M is Einsteinian. Thus from (5) and (6) we see that  $\rho_k^2 = \rho_{k-1} \rho_{k+1}$  and M is Einsteinian. This proves Theorem 1.

If M is cohomologically Einsteinian, then we have

$$c_1 = b\omega$$

for some constant b. Then, by (4), we have

$$a_k = b^k,$$
  $k = 1, \dots, n.$ 

From these we find  $a_k^2 = a_{k-1} a_{k+1}$ . Thus Theorem 2 follows immediately from Theorem 1.

## 3. Proof of Theorem 3

Let M be a complete intersection in  $P_{n+p}(C)$  given as the intersection of p nonsingular hypersurfaces  $M_1, \dots, M_p$ , in general position. Let  $d_a$  denote the degree of  $M_a$ ;  $a = 1, \dots, p$ . Then by a theorem of Riemann-Roch-Hirzebruch ([4, p. 159]) the first Chern class  $c_1$  of M is given by

$$c_1 = \frac{n + p - 1 - \sum d_\alpha}{4 \pi} \omega.$$

This shows that M is cohomologically Einsteinian. Thus, if  $\rho_{k-1}$  and  $\rho_k$  are constants for some k,  $1 \le k \le n$ , then Theorem 2 implies that M in Einsteinian. Thus, by a result of Hano, we see that M is either a linear subspace or a hyperquadric in some (n+1)-dimensional linear subspace. This completes the proof of the theorem.

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